

# Combining Depth Fusion and Photometric Stero for Fine-Detailed 3D Models

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#### Motivation

**Problem**: 3D models obtained from fusion of depth images lack details due to:

- Noise in depth images
- Low resolution
- Smoothing in the fusion process.

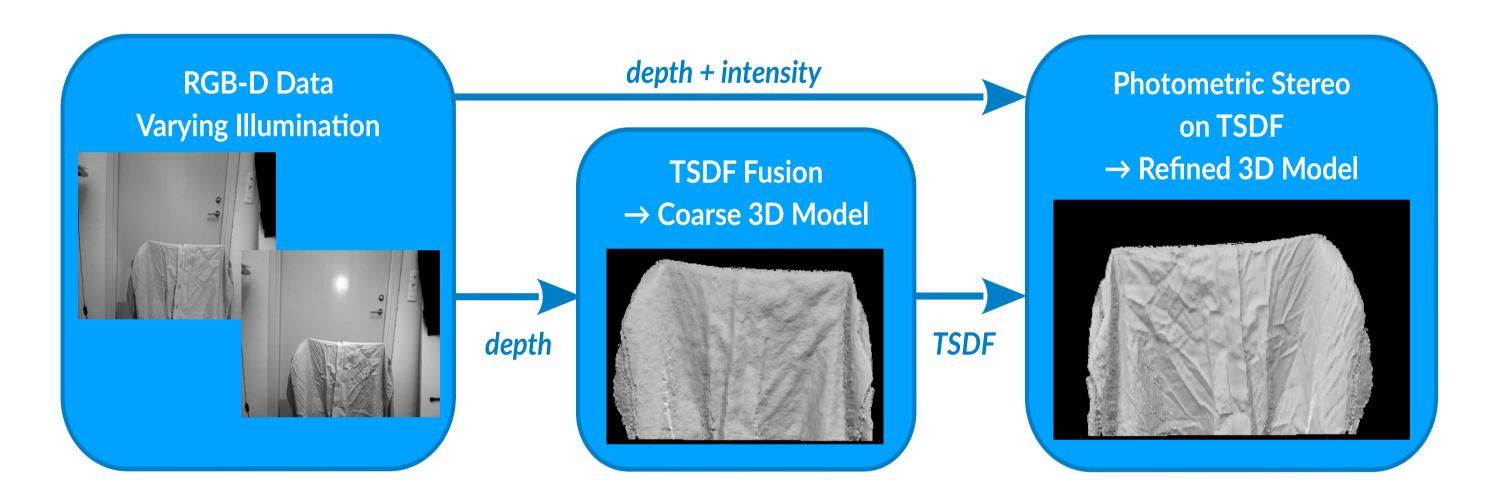
**Goal**: Enhance the quality of the 3D models.

#### Method:

- Capture richer data with varying illumination.
- Combining techniques from Photometric Stereo and Truncated **Signed Distance Functions (TSDF).**

# System Overview

- Input: TSDF, depth- and intensity-images and camera positions
- Reflectance Model: Lambertian model and spherical harmonics
- **Optimiziation:** Optimize over normals, albedo and light sources
- Output: 3D model with more details



#### Notation and Details

 The Lambertian reflectance model estimates the observed intensity in a projected point as

$$\mathcal{I}(\pi(\mathbf{x})) = \rho(\mathbf{x})\mathbf{s}^T\mathbf{n}(\mathbf{x})$$

The normal of a surface point in a TSDF can be computed as

$$\mathbf{n}(\mathbf{x}, \mathbf{d}_V) = \frac{\nabla g_V(\mathbf{x}, \mathbf{d}_V)}{\|\nabla g_V(\mathbf{x}, \mathbf{d}_V)\|}$$

•  $g_V:\mathbb{R}^3 imes\mathbb{R}^8 o\mathbb{R}$  is the tri-linear interpolation function which gives the distance to the surface at point x.

#### Idea:

- Refine the distance values to change the normals in the voxels to better fit the intensity images.
- Captured data with varying illumination contain information about detailed geometry.

#### **Notation**

- $\mathbf{d}_V$  and  $oldsymbol{
  ho}_V$  are eight distance- and albedo-estimates for a voxel V
- $\tilde{\mathbf{s}}, \tilde{\mathbf{n}} \in \mathbb{R}^9$  are the spherical harmonics.
- Surface points are extracted in all voxels that has a zero-crossing,  ${\cal S}$  denotes the set of all such surface points.
- $\mathcal{V}^k$  denotes the set of voxels observed in frame k.

# Optimization

Three error terms are used to improve the 3D model

- Penalize deviation between rendered intensity and observed intensity (1).
- Favor surfaces that are close the observed one in the depth images (2).
- Favor solutions where neighboring voxels have similar albedo (3).

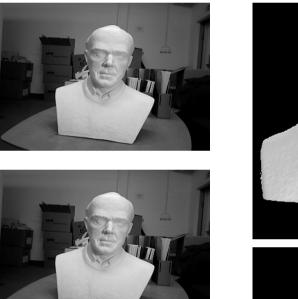
$$E_L(\mathbf{d}, \boldsymbol{\rho}, \tilde{\mathbf{s}}^1, \dots, \tilde{\mathbf{s}}^K) = \sum_{k=1}^K \sum_{V \in \mathcal{V}^k} \sum_{\mathbf{x} \in V \cap \mathcal{S}} (\mathcal{I}^k(\pi(\mathbf{x})) - \rho(\mathbf{x}, \boldsymbol{\rho}_V) \tilde{\mathbf{n}}(\mathbf{x}, \mathbf{d}_V)^T \tilde{\mathbf{s}}^k)^2$$
(1)

$$E_{\text{depth}}(\mathbf{d}) = \sum_{k=1}^{K} \sum_{v \in \mathcal{V}^k} (D^k(\mathbf{x}_v) - d_v)^2$$
(2)

$$E_{\text{albedo}}(\boldsymbol{\rho}) = \sum_{V \in \mathcal{V}} \sum_{v_i \neq v_j \in V} (\rho_{v_i} - \rho_{v_j})^2$$
(3)

## Results

### Shading - Hörmander



Intensity

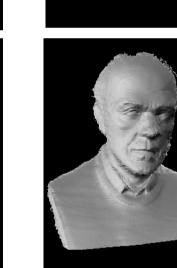
**Images** 

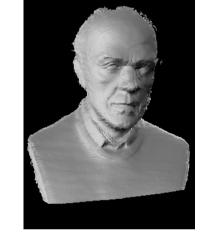




Fused depth

images





Our Result

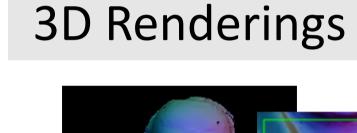


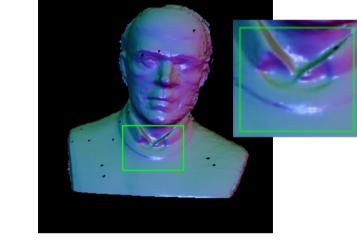


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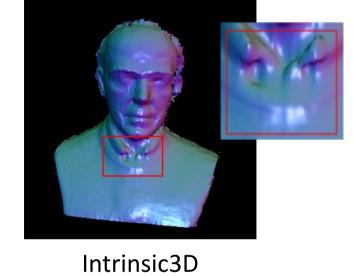


Shading - Shirt

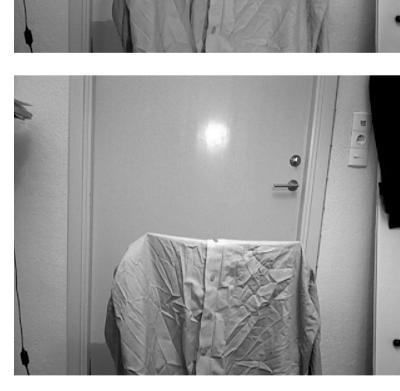




Our result

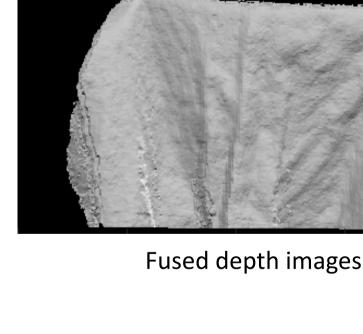


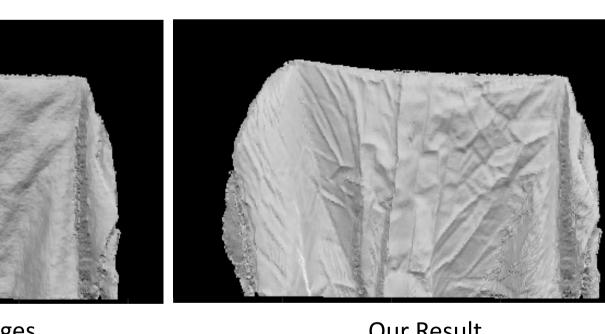
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Intensity Images

 $\mathbf{I}(\pi(\mathbf{x})) = \rho \mathbf{n}^T \mathbf{s}$ 

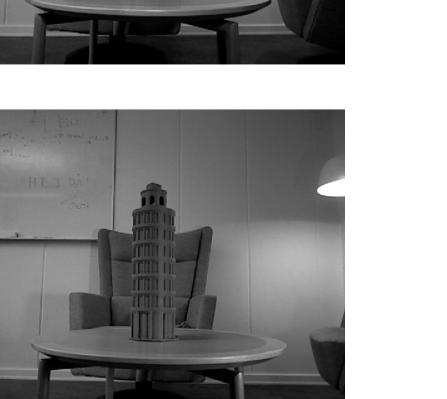




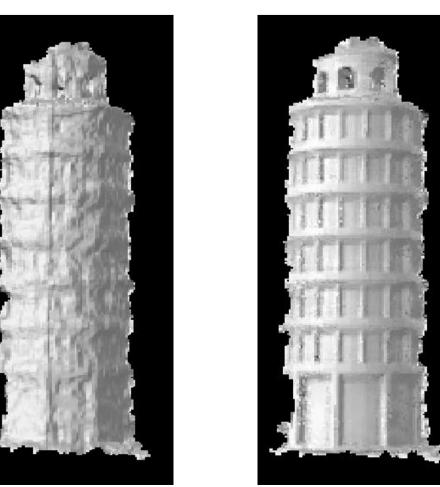
Our Result

Quantitative Result

Compare to groundtruth – 3D printed model



Intensity **Images** 



Fused depth images

Our Result